

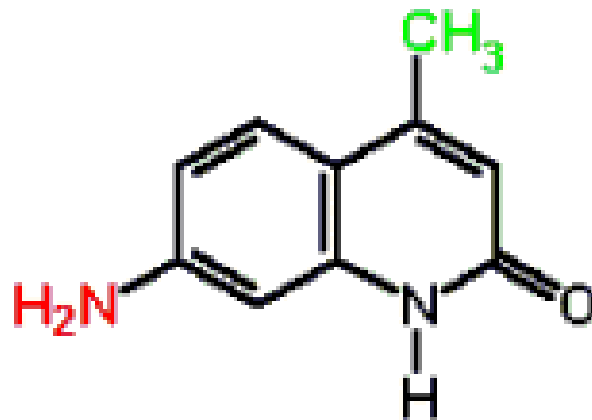
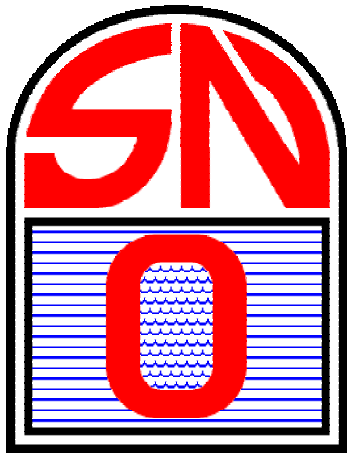
# *ECG Meeting*

*Nov 12, 2003*

## *Carleton University*



### *THE WAVELENGTH SHIFTERS IN SNO A Quick Update*



- SNOMAN
- WLS Telescope
- MC MSW plane

by Etienne Rollin  
Carleton University  
[www.physics.carleton.ca/~erollin](http://www.physics.carleton.ca/~erollin)

# *Two Candidates and SNOMAN*



## *Two Candidates - Two Title Files*

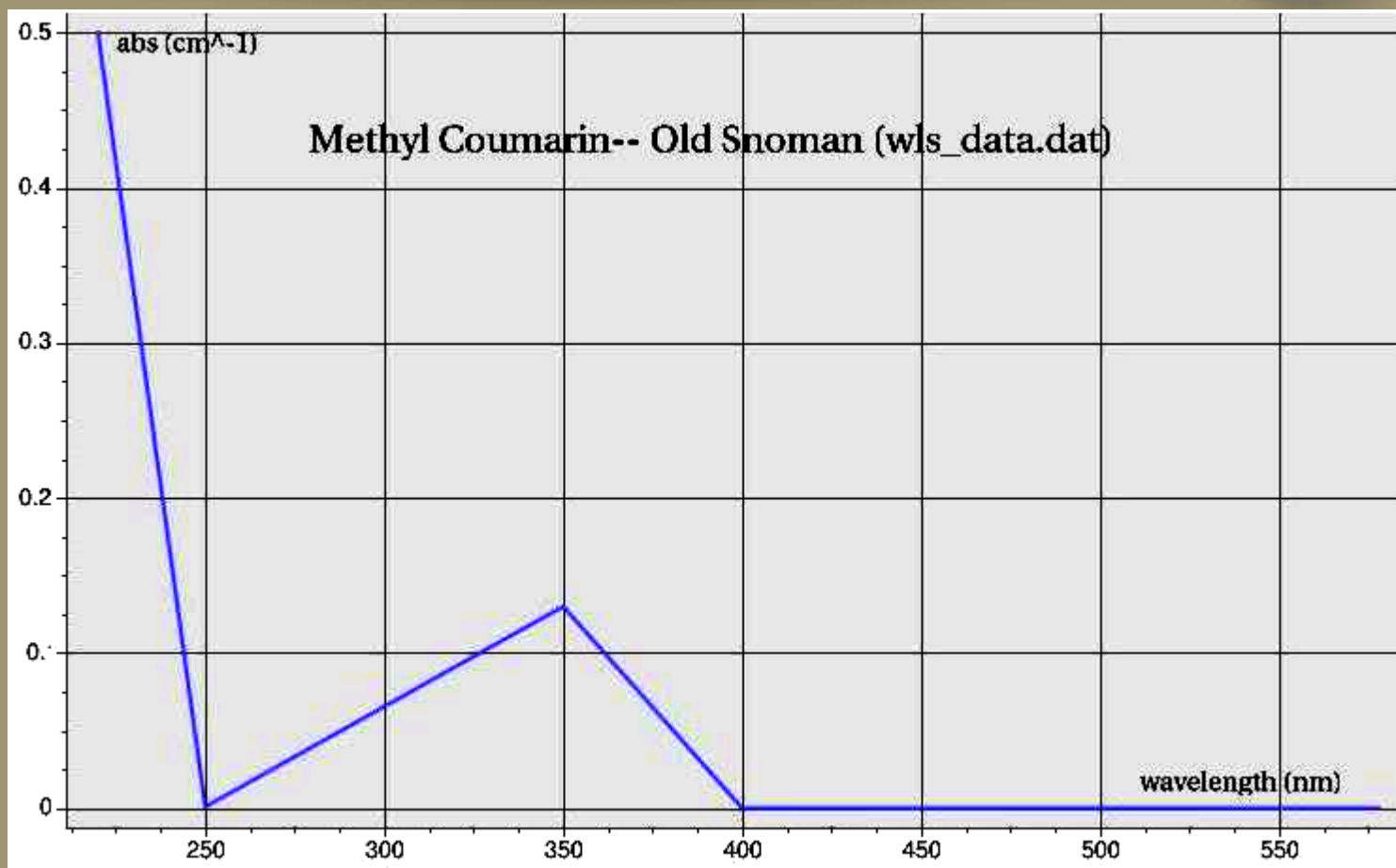
*Carbostyrl 124  
Alexafluor 350*

*\*Title files are available  
on my website*

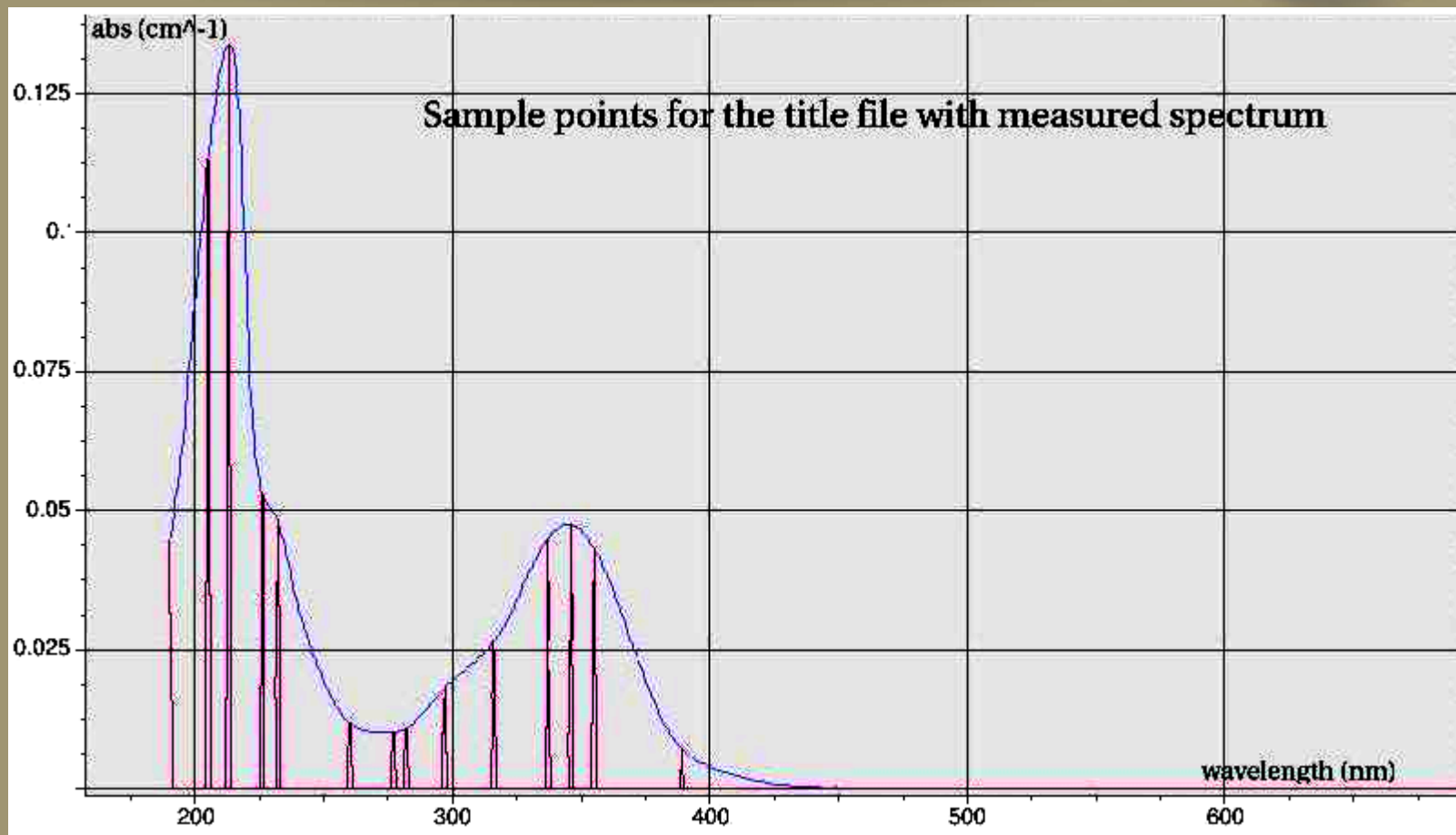
*Before being official:*

- Measurement of the decay time (Nov 19 by Xin)*
- Measurement of the reemission probability (Nov 14)*

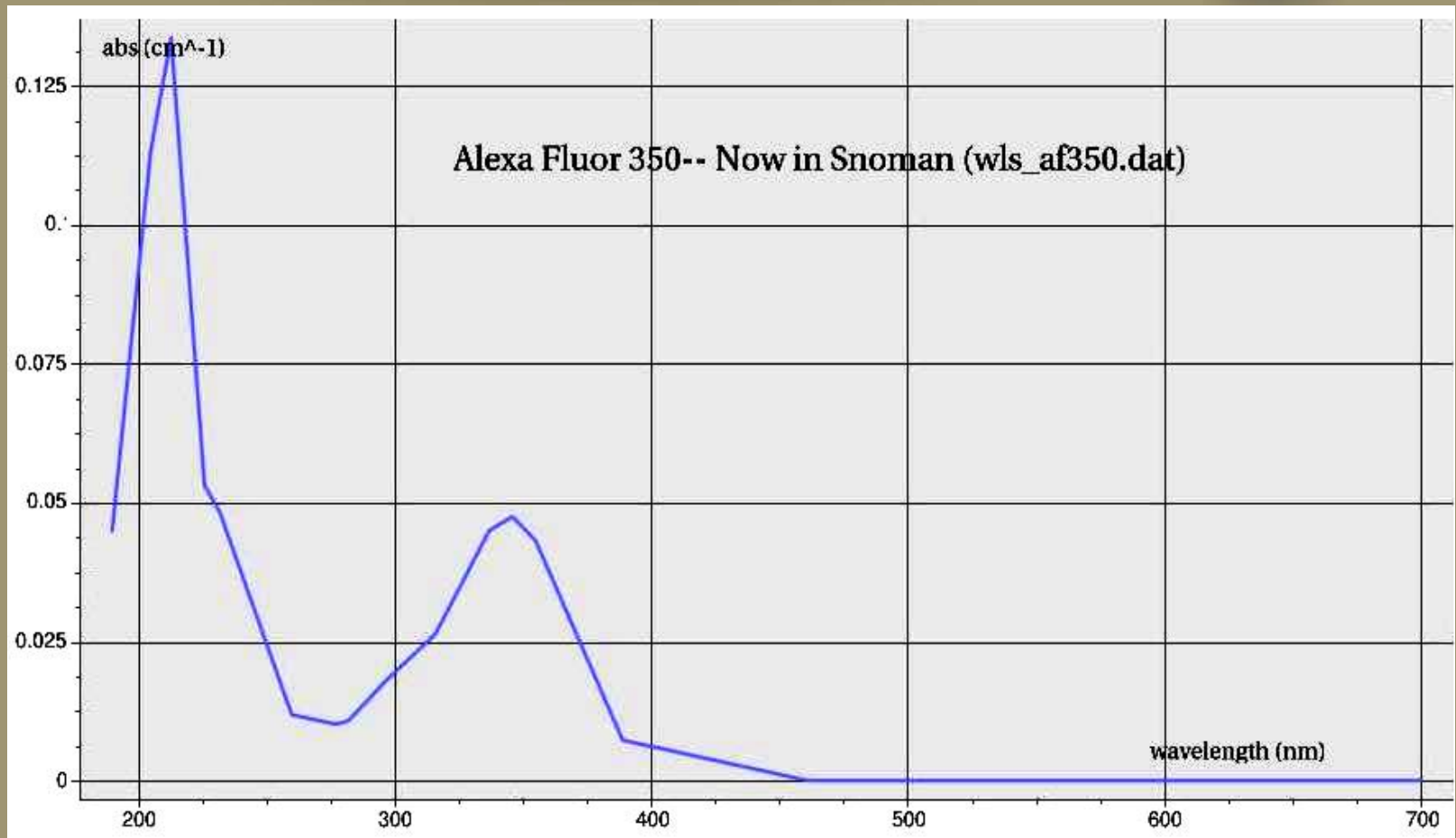
# *Old Titlefile*



# *Measured Spectrum*



*New Titlefile*



# *MC RESULTS!!!*



	<i>Mean Nh̃it for a 10 MeV electron</i>
<i>Pure D<sub>2</sub>O (No Quench)</i>	79
<i>Old Data</i>	148
<i>Alexa Fluor</i>	107
<i>Carbostyryl</i>	199

\*Based on 10000 x 10 MeV e<sup>-</sup> generated isotropically in the center of the detector with a WLS concentration of 1ppm.



# *Cosmic Ray WLS Telescope*



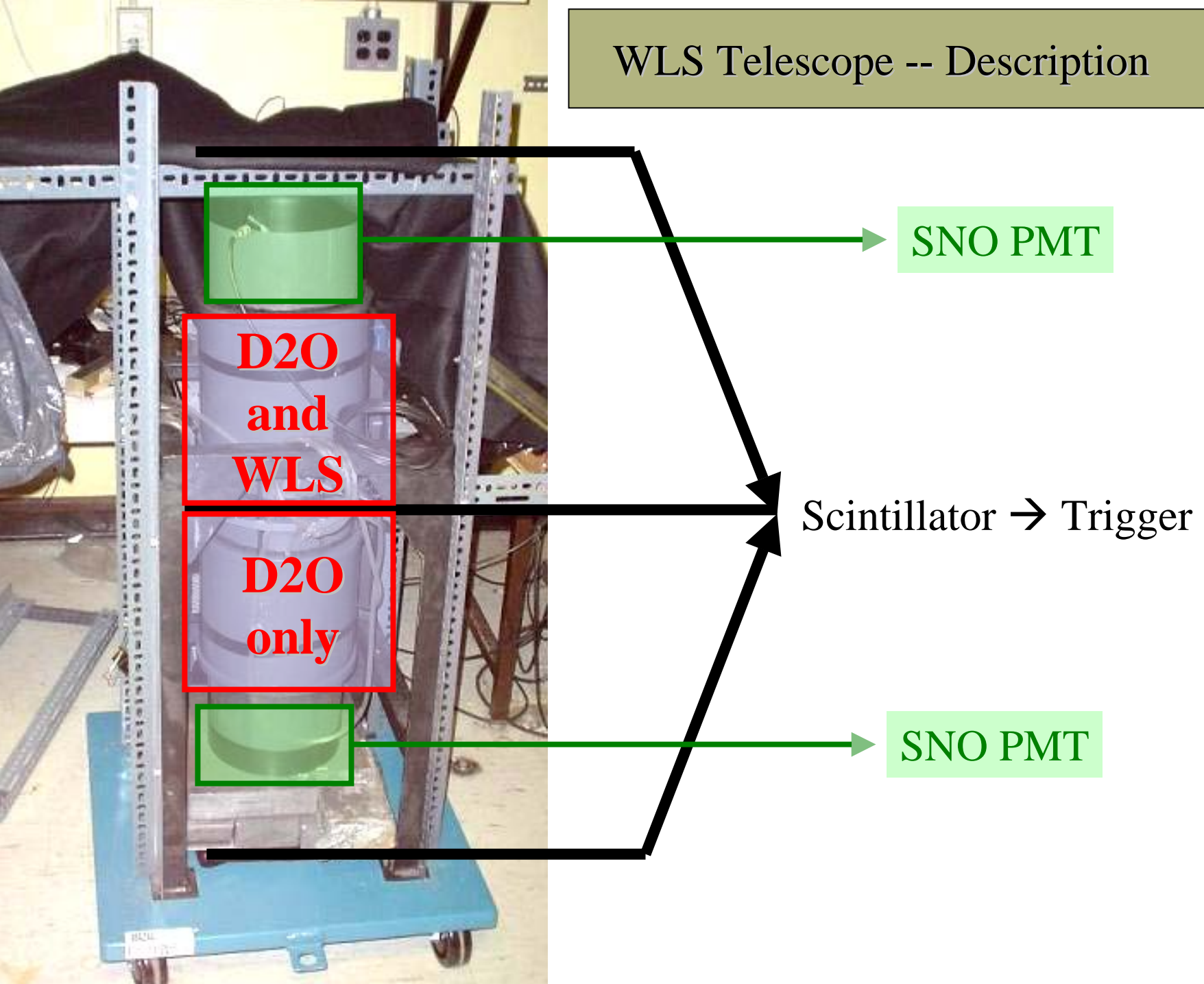
# *Cosmic Ray Telescope*



- *Used to measure properties of WLS*
- *Will help to determine the right concentration to use in SNO*
- *Needs some adjustments to start taking data*



## WLS Telescope -- Description



# *MSW plane*

*With the help of G. Tesic*



*ES + CC + NC + Low E BG*

*(WLS only but may be used with NCD)*

*-Boost the signal*

*Find threshold for the new candidates and fit for new MSW constraints. If WLS only is not meeting expectations (spectral distortions), use the second option.*

# *MSW plane*

*With the help of G. Tesic*



## *ES + CC (WLS+Poison)*

*-Kill all neutrons (no NC or internal background)*

*Both method should measure directly spectrum distortions by comparing undistorted and distorted spectrum (QPhysics) to exclude (?) the hypothesis of undistorted spectrum for the best solution (LMA) or any preferred solutions.*



# Conclusion



*WLS are an interesting phase for SNO because:*

- Increase the signal keeping external bg down*
- Compensate for the aging of the detector*
- Cancel the shadowing effect of the NCD*
- Lower the energy threshold (sensitive to spectral distortions?)*
- Not a major change of the detector (with/without NCD)*
- Maybe: Increase the livetime by decreasing the calibration time since the detector will be more symmetric/uniform (added 20 minutes ago)*